

# Mars Exploration Program (MEP) Update

Presented to the Planetary Protection Subcommittee Dec 8, 2015

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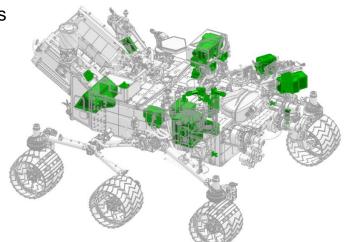


# **MEP Update**

[	oing efforts  ☐ Operational assets continuing to support mission objectives ☐ Mars 2020 development proceeding well, nearing PDR ☐ Many subsystem reviews already complete ☐ MOMA recently delayed due to late delivery of DLR laser
	term ☐ Increasing need to replace and update aging infrastructure ☐ Responding to NRC Decadal Survey science priority of MSR requires additional missions ☐ The 2022 opportunity 5 years from the current budget planning horizon
[	ce/Exploration collaboration  ☐ MEPAG NEX-SAG study complete ☐ Joint MEP/HEOMD Human Landing Site Workshop conducted

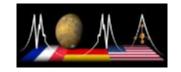
## Mars 2020 Update

- ☐ Conducted 2<sup>nd</sup> Landing Site Workshops in August 2015; ongoing imaging and analysis for top sites
- □ Completed Heritage Flight System Review in September 2015
- Early acquisition and builds of heritage elements and items with low risk of change are proceeding at a fast pace
- Completed Sampling & Caching System (SCS) architecture definition. Working detailed engineering and design for cache system implementation
- Continue working with PPO to finalize the Categorization Letter and PP Plan
- Payload instrument and flight system preliminary design reviews (PDRs) ongoing, culminating in Project PDR in February 2016
  - Completed CEDL, Mastcam-Z, SuperCam, RIMFAX, MEDA, TRN, and Sampling/Caching System PDRs
  - Surface Operations, PIXL, SHERLOC, and MOXIE PDRs scheduled for December-January
- ☐ On track for Feb 2016 PDR



Project has made excellent progress to date

## **MOMA-MS** Project Description



**Project Description:** Delivery of key subsystem to international PI as a part of an ESA Planetary Science mission to explore signs of present/past life on Mars (dual-source mass spectrometer); Class C

### **Science Objectives:**

- MOMA addresses the ExoMars top science goal of seeking signs of past or present life on Mars
- MOMA-MS is a subsystem of MOMA composed of a dual source mass spectrometer to detect a wide range of organic molecules in Martian samples. Organic structure and distribution can be indicators of life

### **Organizational Contributors:**

- ExoMars Mission Lead: European Space Agency
- ExoMars Analytical Lab Drawer: Thales Alenia Space Italia (TAS-I)
- MOMA PI: Max Planck Institute for Solar System Research
  - Gas Chromatograph: LISA/LATMOS (France)
  - Laser: Laser Zentrum Hanover (LZH)
  - Electronics: University of Michigan, Battel Engineering
  - · Wide Range Pump: Creare

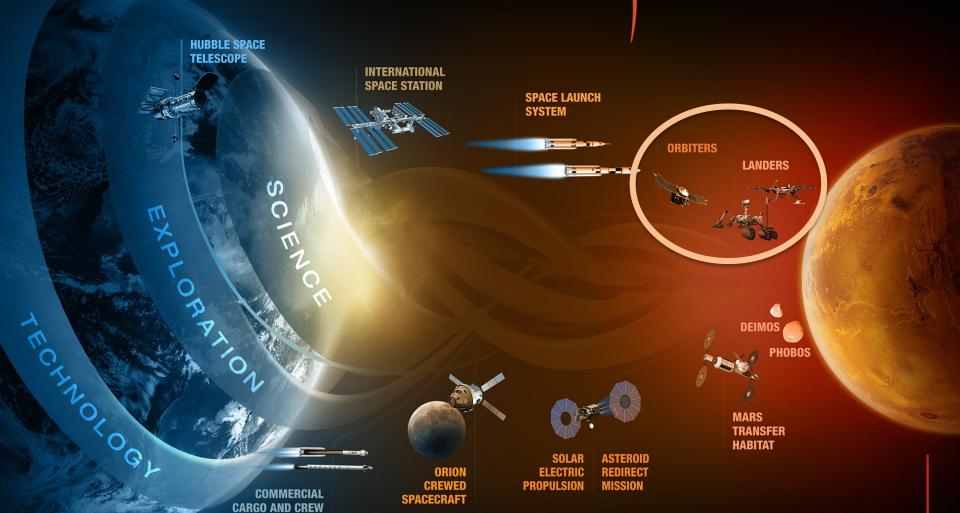
#### **Overall Status:**

- Re-baselined MOMA design to require only LD/MS capabilities due to late delivery of GC from CNES
  - GC option to be kept open as long as possible
- DLR/Max Planck recently informed MOMA team of significant slip in delivery (from Laser Zentrum) and integration (by Max Planck) of flight system laser. Unless resolved, this may push MOMA FM delivery to early 2017 - a major risk to MOMA project ability to meet 2018 ExoMars launch schedule
- Flight MS in LDMS mode (ETU laser)
  - Instrument performance (tested so far) is excellent very clean spectra
- ESA selected Oxia Planum as the landing site for ExoMars 2018



# JOURNEY TO MARS





MISSIONS: 6-12 MONTHS
RETURN: HOURS
EARTH RELIANT

MISSIONS: 1-12 MONTHS RETURN: DAYS MISSIONS: 2-3 YEARS
RETURN: MONTHS
EARTH INDEPENDENT

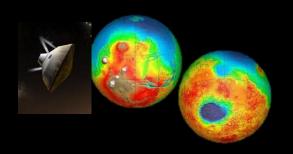
PROVING GROUND

## What We've Learned and Still Need to Learn at Mars

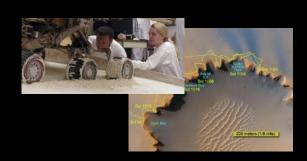
# Orbital environment and operations



# Capture, EDL & Ascent at Mars



### Surface Operations at Mars



#### Learned:

Deep space navigation

Orbit transfer near low-gravity bodies

**Gravity assist** 

Aero-braking

**Gravitational** potential

Mars' moons characteristics

ISRU potential

#### To Learn:

Return flight from Mars to Earth Autonomous Rendezvous & Docking

ISRU feasibility

Resource characterization of Mars moons

High-power SEP

#### Learned

Spatial/temporal temperature variability

Density and composition variability

Storm structure, duration and intensity

1 mT Payload

~10 km Accuracy

#### To Learn

Ascent from Mars

Large mass EDL

**Precision EDL** 

Aero-capture

Site topography and roughness

Long-term atmospheric variability

#### Learned

Water once flowed and was stable

Global topography: elevation and boulder distributions

Remnant magnetic field

Dust impacts on Solar Power / Mechanisms

Radiation dose

Global resource distribution

Relay strategies, operations cadence

#### To Learn

Landing site resource survey

Dust effects on human health, suits & seals

Rad/ECLSS in Mars in environment

Power sufficient for ISRU

**Surface Navigation** 

Resource prospecting and round trip experience are key enablers

# Conceptual Integrated Campaign for Mars Precursors "in the 2020s"

Mars 2020

Mars Orbiter

Round-Trip Surface to Surface

Exploration Precursors



**LEGEND** 

Exploration

Cross-Cutting (Exploration-Technology-Science)

Science

ISRU Prototype

EDL Instruments

Sample Acquisition

> In Situ Science

Habitable Conditions

Ancient Life Resource Survey

Landing Site Selection

Optical Comm/Relay

**SEP** 

Rendezvous

Remote Sensing Instruments **Dust Toxicity** 

EDL Evolution/ Instruments

Mars Ascent

Surface Navigation

Returned Sample Analysis ISRU Production

Surface Power for ISRU

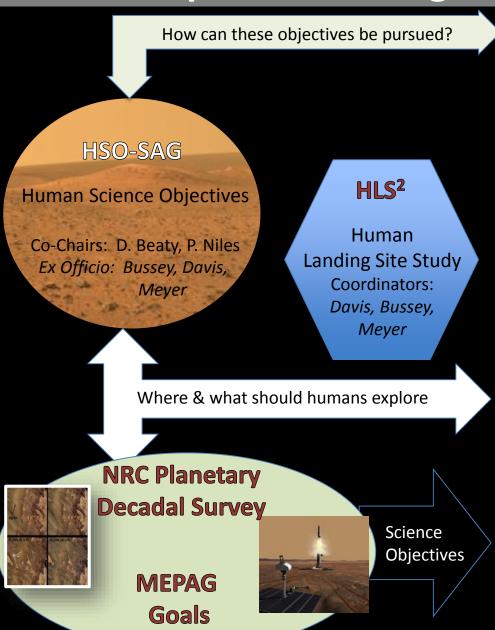
Rad/ECLSS

Validation

Increased EDL Mass & Precision

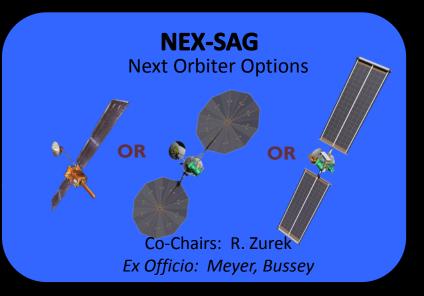
Science Instruments

# **Science Exploration Integration**





What are the Base & Exploration Zone criteria? What & where are the resources needed?

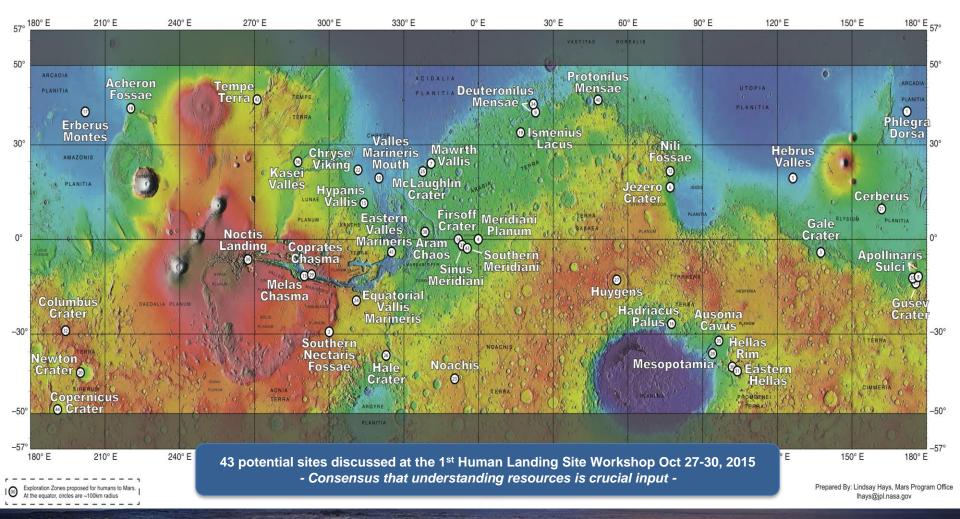


## **Next Orbiter (NEX-SAG) Findings**

- SEP brings the advantages of orbit flexibility and increased payload mass & power
- Advanced telecom provides necessary coverage for high-resolution data
- Considerable overlap between science goals and human exploration resource prospecting interests & derived objectives yield similar, mature instrument approaches
  - Visible imaging of HiRISE-class or better (~15-30 cm/pixel)
  - Polarimetric synthetic aperture radar imaging with penetration depth of a few (<10)
    meters and spatial resolution of ~15 m/pixel to search for shallow ground ice and crustal
    structure</li>
  - Short-wave IR spectral mapping with a spatial resolution of ~6 m/pixel (3 x CRISM) with sufficient spectral resolution to detect key minerals
  - Long-wave atmospheric sounding for wind, temperature, & water vapor profiles with 5 km vertical resolution
  - Thermal IR sounding for aerosol (dust & ice) profiles
  - Multi-band thermal IR mapping of thermo-physical surface properties (e.g., ice overburden) and surface composition
  - Wide-angle imaging to monitor weather and surface frosts (global, km-scale)



#### **Potential Exploration Zones for Human Missions to the Surface of Mars**





## Mars Exploration Program Status Summary

